



# Latest developments of adaptive optics for ultra intense lasers and plasma diagnostics.

Xavier Levecq Target to Plasma 2015



## **Outline**

- Adaptive optics introduction
- Full correction strategy
  - Standard adaptive optics principle
  - Phase retrieval adaptive optics
- Plasma diagnostics
  - Wavefront diagnostics
  - Example of dense plasma diagnosis with EUV light source

- Conclusion



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#### Adaptive optics introduction



For lasers applications, the wavefront distortions are caused by: - very high number optical elements

- misalignment of optical elements (beam expanders, ...)

- Thermal effects in the amplifiers

- mechanical constraints on large optical elements

These wavefront distortions degrade the focused beam and decrease the maximum focused intensity





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#### Standard adaptive optics principle





#### Full correction adaptive optics strategy





Experts in adaptive optics for ultra intense lasers

#### Full correction adaptive optics strategy : step 1, acquisition





Experts in adaptive optics for ultra intense lasers

1.00 2.00

3.06

Acquired images

#### Full correction adaptive optics strategy : step 1, acquisition





um

2 00

**Acquired images** 

-2.00 -1.00 0.00 1.00

#### Full correction adaptive optics strategy : step 1, acquisition





#### Full correction adaptive optics strategy : step 2 , calculation





#### Full correction adaptive optics strategy : step 3, correction





#### Full correction adaptive optics strategy : step 3, correction





#### Near Future : new AO strategies dedicated to laser users

In the near future, a single laser will be used for multiple experiment areas.





#### Adaptive optics at Imagine Optic

# Imagine Optic provides and manufacture complete adaptive optics solutions

#### WAVEFRONT SENSORS

- EUV (6 to 60 nm), UV (193-266 nm), VISIBLE, NIR (600-1100nm)
- X (up to 25 keV) synchrotron applications





#### **DEFORMABLE MIRRORS**

- *MIRAO 52-e* : *low power, magnetic mirror: microscopy applications*
- ILAO : high power, mechanical mirror: laser application

#### **SOFTWARE**

- Wave View : Metrology tools
- Wave Tune : usual « close loop » adaptive optics
- PhaRAO : specific adaptive optics
- WaveTune : AO software with focal spot camera





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#### **Plasma Diagnostic : interferometry or wavefront sensor**



G.R. Plateau et al, Rev. Sci. Instrum. 81, 033108 2010



Plasma Diagnostic : which wavelength?

- Sub critical density < 10<sup>20</sup> e-/cm<sup>3</sup> Wavelength : UV or visible

- Dense plasma between 10<sup>20</sup> and 10<sup>20</sup> e-/cm<sup>3</sup> Wavelength : EUV (5 – 100 nm)

- Overdense plasma > 10<sup>23</sup> e-/cm<sup>3</sup> Wavelength : X-ray (5 – 25 keV)







#### Plasma Diagnostic : Dense plasma with EUV





#### **Plasma Diagnostic : Dense plasma with EUV**



Ongoing data analysis:

- Re-tracing HHG propagation to know spot profile at plasma



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### Conclusions

Imagine Optic is strongly committed in the development of adaptive optics solutions for High Power Laser applications:

• ILAO mechanical deformable mirror has been developped to respond to specific needs of adaptive optics for demanding laser applications Coatings, replacable substrate, Diameters 50 to 335mm, High quality, stable correction

• New adaptive optics correction strategies are being developed for High Power Laser users

**Phase Retrieval adaptive optics :** Direct (no iterations) and fully automated method (no complex user operation)

- Beam diagnostic systems,
- Plasma diagnostic tools
- Complete support and consulting
- and more to come in the future ...



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## Thank you for your attention

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## **Achromaticity of Shack Hartmann WFS**

## SH wavefront sensor is strictly achromatic



When the wavelength changes, the spots are not at the perfect focus on the CCD. **However, the spot position remains the same** (What matters is not the focal lenght of the microlenses but the distance between the microlenses and CCD).

#### Same spot postion $\rightarrow$ same WF measurement = SH wavefront sensor is achromatic.

SH is used in astronomy on sources with blackbody spectrum

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#### Adaptive optics developed for astronomy

Atmospheric turbulence creates wavefront distorsion.

Wavefront distorsion decreases image resolution in ground-based telescopes.





#### **Deformable mirror : mechanical actuators technology**

Principle : generation of a force at the back of a substrate





→ Conversion of macroscopic displacements in microscopic shape change





Integrated with High Power dielectric coating from state of the art manufacturer

52 actuators



#### **Deformable mirror : Mechanical actuators technology**

Features :

- Number of actuators : ~10 -> ~100 (between 30 55 typical)
- Stroke : 100  $\mu m$
- Size : 30 600 mm in diameter (no technological higher limit)
- Optical quality : 10 20 nm rms WFE
- Compatible with any coating : Metal, Hybrid, Dielectric (usual manufacturers

Advantage :

- Scalability
- Stability (even unpowered)
- EMC immunity
- Simple maintenance (actuators and reflective substrate)

Drawbacks :

- Speed (<0.5Hz)
- Backlash compensation









## Outline

- Adaptive optics introduction
- Overview of the state of the art of adaptive optics components
  - Deformable mirror
  - Wavefront sensor
- Perspectives for near future and next generation laser facilities
  - Software developments : use of a focal spot camera
  - Software development : Tango and Epics control
  - Hardware development : Beam reducer
- Conclusion



## Overview of the adaptive optics

## Components: wavefront sensor Imagine Optics manufactures Shack Hartmann wavefront sensors

Shack Hartmann wavefront sensors are:

- achromatic
- insensitive to vibrations
- large dynamic range
- measure wavefront everywhere there are photons (no assumption on laser beam pupil shape)

Shack Hartmann : everywhere there are photons  $\rightarrow$  Spots on the CCD  $\rightarrow$  phase and intensity is measured

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## **Typical laser beam measurement**



Wavefront and intensity are measured even at the extreme edge of the beam.

- •Independant of beam shape
- •Whole beam taken into account



## Near Future: total integration of the

# AO in the control command of the



area #3

→ A big work has to be done on in order to have a AO system that can be fully and easily controlled through TANGO or EPICS systems.